Proton elastic scattering from ⁷Be at low energies.

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The elastic scattering of protons on ⁷Be has been measured in the energy region from 1-3.3 MeV via the thick-target technique. The data conclusively demonstrate the existence of a 2⁻ state at an excitation energy of approximately 3.5 MeV in ⁸B, and rule out a predicted 1⁺ state near 1.4 MeV. The relevance of these results for the ⁷Be (p, γ) reaction, of interest in solar neutrino physics, is discussed.

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The ${}^{7}\text{Be}(p,\gamma){}^{8}\text{B}$ reaction plays a very important role in nuclear astrophysics, since it is the source of the highenergy solar neutrinos detected in the Cl neutrino experiment [1]. There have been many measurements of this critical reaction, both direct using a radioactive ⁷Be target (see Ref. [2] for a comprehensive review of these experiments), and indirect via Coulomb dissociation [3–5] of ⁸B. The astrophysical S-factor [6] deduced from these experiments is dominated by the narrow 1^+ resonance at 0.63 MeV which plays no role at solar energies (20 keV). However, it is not possible to measure the cross section at 20 keV so one must rely on extrapolations from higherenergy data. After evaluating all the direct measurements, Adelberger, et al. [2] give a recommended value of $S_{17} = 19^{+4}_{-2}$ eV b at zero energy (1 σ error), and state that further measurements are desirable to reduce the uncertainties below 5% in order to achieve a full understanding of the data from new solar neutrino experiments.

The present work was occasioned by recent experimental [7] and theoretical [8] work which relates to the extrapolation of the (p,γ) data to solar energies. Gol'dberg, et al. [7] report evidence for a broad, s-wave level of ${}^{8}B$ at an excitation energy of approximately 3 MeV from a study of proton elastic scattering on ⁷Be. They suggest 1^{-} or 2^{-} for the spin/parity of this level but were unable to give a good account of the width of the resonance due to the rather poor statistics in the experiment. Csótó [8] used a microscopic cluster model to predict the existence of a second 1^+ state at low excitation energy which is also broad enough to have important consequences for the S(0) value. We have measured proton elastic scattering from ⁷Be in the ⁸B excitation-energy range from 1-3.3 MeV in order to clarify the resonance structure in this important region.



FIG. 1. Time-of-flight vs. energy for the secondary beam, showing the major contaminant groups. Note that 7 Be is easily separated by time of flight.

The experiment was carried out using the *TwinSol* radioactive ion beam (RIB) facility [9]. A 2.5 cm long gas target containing 1 atm of ³He was bombarded by a highintensity (up to 100 particle-nA), nanosecond-bunched primary ¹⁰B beam at an energy of 51.0 MeV. The entrance and exit windows of the gas cell consisted of 2.0 μ m Havar foils. The secondary beam was momentum selected and transported through the first of two superconducting solenoids, which focused it onto an 8.0 μ m

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Havar foil. Differential energy loss in this foil allowed for the purification of the beam as it passed through the second solenoid. The laboratory energy of the ⁷Be beam at the secondary target position was 25.5 MeV, with a resolution of 1.5 MeV full width at half maximum (FWHM) and an intensity of up to 1.0×10^5 particles per second. The energy spread was due to a combination of the kinematic shift in the production reaction and energy-loss straggling in the gas-cell windows and the energy-loss foil. The beam had a maximum angular divergence of $\pm 4^{\circ}$ and a spot size of 5 mm FWHM. Although contaminant ions were still present in the beam, they could be identified using time-of-flight (TOF) techniques. The TOF of the particles was obtained from the time difference between the occurrence of an E signal in a detector telescope and the RF timing pulse from the beam buncher. The time resolution of better than 3 ns (FWHM) was adequate to cleanly separate ⁷Be from all other ions. This is illustrated in Fig. 1, which was obtained with a Si ΔE -E telescope placed directly in the secondary beam, after reducing the primary beam intensity by three orders of magnitude. Some of the more important contaminants are indicated. The intensity of the beam during the experiment was determined by relating the number of ⁷Be ions in this figure, to the integrated charge of the primary 10 B beam collected in the *TwinSol* Faraday cup.

The ⁷Be beam was stopped in a 12.0 mg/cm² thick CH₂ target. The recoil protons from back-angle elastic scattering in this target lose only a small amount of energy in traversing the foil, and emerge from it with sufficient energy to be detected. Note that the lowest-energy protons, from the scattering of ⁷Be ions near the end of their range, encounter the least amount of material before leaving the target. In this way, an excitation function for elastic scattering down to very low energies can be measured with high efficiency and good resolution. The recoil protons were detected with two telescopes consisting of 17.8 and 19.2 μ m Si ΔE detectors, backed by 1000 μ m Si E detectors. The active area of the ΔE detectors was 450 mm^2 , and that of the E detectors was 600 mm^2 . Each telescope had a circular collimator that subtended a solid angle of 11 msr. They were placed on either side of the beam at $\Theta_{\text{LAB}} = \pm 15^{\circ}$. It would have been preferable to place a telescope at 0° to the beam, but the light-ion contamination (Fig.1) produced a count rate in this position that was unacceptable since these ions penetrated the target and directly entered the telescope. It was verified that the recoil proton TOF signal was only slightly shifted in time relative to ⁷Be and was stable during the course of the experiment, so that the separation from elastic scattering of contaminant ions was excellent. Lowenergy protons can also result from fusion of ⁷Be with C in the target. We measured a spectrum for this process using a thick C target, and subtracted it from that obtained with the CH_2 target. Finally, inelastic scattering of ⁷Be on ¹H can occur at energies high enough to popu-



FIG. 2. Excitation function for ${}^{7}\text{Be+p}$ elastic scattering at a cm angle of $148 \pm 12^{\circ}$ measured in the present experiment. The data are compared with an R-matrix calculation including the well-known states 1^{+} and 3^{+} states at 0.77 MeV and 2.32 MeV, respectively. The dashed curve is the pure R-matrix prediction, while the solid curve is the R-matrix calculation convoluted with the experimental resolution function. See text for a discussion of this calculation.

late the first excited state of ⁷Be. However, as discussed below, the cross section for this process is negligible and it will be ignored here.

The cross section for back-angle proton scattering from ⁷Be measured in this experiment is shown in Fig. 2, compared with an R-matrix calculation that includes nonresonant, direct scattering and the well-known first 1^+ and 3^+ states [10] in ⁸B. Since we have a limited set of experimental data consisting of an excitation function for elastic scattering at only one angle, additional information was used in order to obtain a reliable answer from the R-matrix fit. The goal was to fix as many of the parameters in the R-matrix calculation as possible. A reasonable approach is to use the mirror ⁷Li+n system as a reference point. A large amount of experimental data, along with a comprehensive R-matrix analysis, is available for this system (see Ref. [11] and references therein). Also, the two well known states in ${}^{8}B$ in the measured excitation energy range have well-defined resonance parameters [10]. Using the R-matrix parameters, such as the channel radius, from Ref. [11], and the excitation energy and width of the two resonances in ⁸B, one can obtain a prediction for the excitation function with no free parameters, shown as the dashed curve in Fig. 2. The background phase shift was determined from the Coulomb phase shift plus the phase shift for hardsphere scattering calculated at a channel radius a=4.3 fm [11]. A one-channel approximation was used since the elastic-channel penetrability factors for both the 1⁺ and 3⁺ states far exceed that of all other channels. (The 1⁺ state is not seen in this figure because its excitation energy is out of the measured range).

The solid curve in Fig. 2 shows the R-matrix calculation convoluted with an energy-dependent resolution function which includes the effect of energy straggling of ⁷Be and ¹H in the target, the kinematic shift over the aperture of the detector telescope, and the spot size, energy resolution, and angular divergence of the incident beam. This resolution function was computed with Monte Carlo techniques from the known parameters of the experiment. (Note that the dropoff in the measured cross section at energies above 3 MeV is a direct result of the finite energy resolution of the ⁷Be beam). The absolute normalization of the data was determined in the experiment and has not been adjusted to improve the fit. It can be seen that, as in the previous work [7], there is clear evidence for a very large amount of additional resonance strength that is not accounted for by the levels shown in Ref. [10]. Our data are consistent with those of Ref. [7], but the much-improved statistics allow for more definitive conclusions regarding the parameters of these "missing" state(s).

The introduction of a 2^- (2s) state having width greater then 4 MeV and energy about 3.5 MeV produces very good agreement with the data, as shown by the solid curve in Fig. 3. The structure at 2-3 MeV is now the result of interference between the 3^+ and 2^- resonances. It is important to note that again there has been no adjustment of the background phase shift, which was fixed and equal to the hard sphere value, nor of the absolute cross section. The only two free parameters are the energy and width of the 2^- state.

Attempts to assign spin-parity 1^- to this new resonance failed due to the fact that a 1^- state has channel spin s=1 and does not interfere with the 3^+ state. The elastic-scattering cross section of the presumed 1⁻ state itself is insufficient to reproduce the observed cross section in the $E_x=2-3$ MeV energy region. We also failed to obtain a good fit with spin-parity 1^+ for the resonance in question since the cross section was too low even though a 1^+ state could have channel spin s=2 and therefore interfere with the 3^+ state. The typical interference pattern for these two states results in a minimum in the cross section, which in our case should be near 2.5 MeV and is not observed in the experimental data. (We have assumed that the 1^+ and 1^- states are pure single particle resonances with only one decay mode - via the elastic channel. The assumption that other modes of decay are present would lead to even lower cross sections and worse agreement with the experimental result). Still another possibility is spin-parity 2^+ for the resonance. Two arguments are against this: a 2^+ state with noticeable elastic



FIG. 3. Elastic data compared with the full R-matrix calculation, including the 2^- level at 3.5 MeV. The dashed curve shows the calculation prior to convolution with the experimental resolution function. The dotted curve shows the comparison with a calculation including a predicted 1^+ state at 1.4 MeV. The insert gives a confidence band for the excitation energy and width of the 2^- resonance. The cross on the insert indicates the best fit point, and the solid line is the Wigner limit for the 2^- state calculated for a channel radius of 4.3 fm.

partial width would be a strong resonance which should be observed in the elastic scattering of neutrons from ⁷Li. But the closest 2^+ state in ⁸Li is found only at 4.76 MeV [11], so it is highly improbable that this resonance occurs as low as 3 MeV in ⁸B. Also, the interference of 3^+ and 2^+ states again leads to a minimum in the cross section between resonances, contrary to experiment.

We have not included an R_0 term [12,13] in our analysis. This purely phenomenological term is often used to model the influence of unknown high lying resonances. However, the density of states in A=8 nuclei is very low at the excitation energies of interest. For example, the next 2^- and 3^+ resonances are observed at excitation energies near 7 MeV in ⁸Li [11]. Accounting for these resonances had only a negligibly small influence on the elastic cross section at our energies. Thus, neglecting high-lying resonances seems to be a very good approximation in this particular case, and it has the virtue that additional, arbitrary parameters are not introduced into the fitting procedure. Note also that the resonance spectrum is dominated by the 2^- and 3^+ states, which could decay to the first excited state of ⁷Be via angular momentum $\ell = 2$ and 3, respectively. However, since the penetrability factors for such large ℓ values are very small, inelastic decay can be ignored.

The insert in Fig. 3 shows a confidence band for the energy and width of the 2^- resonance. This state is very broad, which prevents us from obtaining a really precise determination of its parameters. The excitation energy and width extracted from the R-matrix analysis are correlated, as expected. The best-fit point, shown on the figure as a cross, corresponds to $E = 3.5 \pm 0.5$ MeV, $\Gamma = 8 \pm 4$ MeV. The lower limit on the width is well defined, but the analysis is less sensitive to the upper limit. We therefore show the Wigner limit for the singleparticle width (calculated for a channel radius a = 4.3fm), plotted as the solid line in Fig. 3. It can be seen that the 2^- state most likely has a width that is close to the Wigner limit. This observation is in agreement with the data on elastic scattering of neutrons on 7 Li [11]. A broad 2^{-} state (4 MeV or greater) at an excitation energy of about 3.2 MeV is required to fit the a_2 scattering length for $^{7}Li+n$ [14].

Barker and Mukhamedzhanov [14] have analyzed the effect of a 2^{-} state with parameters reported by Gol'dberg, et al. [7] on the astrophysical S_{17} factor at low energies. They noted that previous calculations of the Sfactor have *implicitly* included such a state through its effect on the low-energy $^{7}Li + n$ elastic-scattering phase shifts. However, given the excitation energy $(E_x = 3.0)$ MeV) and width ($\Gamma^0 = 1-4$ MeV) reported in Ref. [7], a channel radius a = 4.0 fm is required in the R-matrix calculation of the ${}^{7}\text{Be}(p,\gamma){}^{8}\text{B}$ S factor. This leads to $S_{17}(0)$ = 16 ev b and an excitation function which lies below essentially all the capture data at low energies. The resonance parameters determined in the current work will result in $S_{17}(0)$ closer to 18 ev b. The predicted excitation function above $E_{cm} = 1$ MeV will then be consistent with that of Kikuchi, et al. [4].

In addition to the negative parity state, there was an indication of a 1^+ resonance at an excitation energy of 2.8 MeV reported in Ref. [7]. We have found that it is not necessary to include such a state to obtain very good fits to our experimental data. On the other hand, we cannot rule out this resonance, especially if it has large reduced width for decay into the inelastic channel. A 1^+ resonance in ⁸Li was observed at a slightly higher energy (3.46 MeV) and it has large reduced width for inelastic decay [11]. The penetrability factors for elastic and inelastic decay of the 1^+ resonance are almost equal, so a resonance with these properties would be hard to detect in our measurements on top of the "background" from the interference of the 2^- and 3^+ states. The same is true of the analog of the 1^- resonance at energy 3.47 MeV in ⁸Li [11].

Yet another 1^+ resonance in ⁸B has been proposed by Csótó [8] in recent theoretical work. Using a microscopic three-cluster approach, he predicted the existence of a 1^+ state at an excitation energy of only 1.4 MeV, with large elastic and negligible inelastic reduced width, and a calculated total width of 560 keV. A state with these parameters should be readily apparent in our excitation function, if it exists. We observe no indication of it in the present experiment. To demonstrate the sensitivity of our measurements, we added a 1^+ resonance with the properties predicted by Csótó into the R-matrix calculation. The result is shown as the dotted curve in Fig. 3. The ratio of elastic to inelastic reduced width would have to be at least 50 times smaller than predicted by Csótó in order to make this calculation agree with experiment.

In conclusion, we have measured back-angle elastic scattering of protons on ⁷Be in the region from 1-3.3 MeV via the thick-target technique. A predicted broad 1^+ state at $E_x = 1.4$ MeV, which would have very important consequences for $S_{17}(0)$, was shown not to exist. There is no evidence for any other 1^+ state in the energy region investigated. A state at $E_x = 3.5$ MeV has been shown to have spin/parity 2^- . Its location and width have been determined with far better accuracy than in a previous experiment [7]. This state has implicitly been included in the calculations of S_{17} , but its properties determined in the present experiment will place further constraints on the theoretical calculations and thus improve the precision of the S-factor at solar energies.

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- [1] R. Davis, Jr., Prog. Part. Nucl. Phys. 32, 13 (1994).
- [2] E.G. Adelberger et al., Rev. Mod. Phys. 70, 1265 (1998).
- [3] T. Motobayashi, et al., Phys. Rev. Lett. 73, 2680 (1994).
- [4] T. Kikuchi, et al., Phys. Lett. B391, 261 (1997).
- [5] N. Iwasa, et al., Phys. Rev. Lett. 83, 2910 (1999).
- [6] J.N. Bahcall, *Neutrino Astrophysics* (Cambridge University Press, Cambridge, England, 1989).
- [7] V.Z. Gol'dberg, G.V. Rogachev, M.S. Golovkov, V.I. Dukhanov, I.N. Serikov, and V.A. Timofeev, JETP Lett. 67, 1013 (1998).
- [8] Atilla Csótó, Phys. Rev. C 61, 024311 (2000).
- [9] M. Y. Lee, *et al.*, Nucl. Instrum. Methods in Phys. Research A422, 536 (1999).
- [10] F. Ajzenberg-Selove, Nucl. Phys. A490, 1 (1988).
- [11] H.D. Knox, D.A. Resler, and R.O. Lane, Nucl. Phys. A466, 245 (1987).
- [12] A.M. Lane and R.G. Thomas, Rev. Mod. Phys. 30, 257 (1958).
- [13] H.D. Knox and R.O. Lane, Nucl. Phys. A359, 131 (1981).
- [14] F. C. Barker and A.M. Mukhamedzhanov, Nucl. Phys. A673, 526 (2000).